

# A STUDY OF INTER-SPACECRAFT COULOMB FORCES AND IMPLICATIONS FOR FORMATION FLYING

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## ABSTRACT

In the course of exploiting spacecraft formations for use in separated interferometry (or other missions), it is possible that the separation distance between vehicles will be on the order of ten meters. This paper investigates the effects of spacecraft charging on the dynamics of very closely spaced formations. For certain high Earth orbits, the ambient plasma conditions will conspire to produce significant spacecraft charging in an environment with a plasma Debye length of more than 100 m. For such conditions, this paper shows the potential to develop disruptive inter-spacecraft Coulomb forces and torques with magnitude comparable to candidate formation-keeping thrusters over distances of tens of meters. Owing to the unexpectedly large interaction forces, the paper also explores the concept of purposely charging spacecraft to affect formation-keeping Coulomb forces. Analytic methods are developed that show the existence of static equilibrium formations in Earth orbit using only inter-vehicle Coulomb forces for one-, two- and three-dimensional formations. Such Coulomb formations would be free of the risk of plume contamination due to thrusters firing in close proximity. Figures of merit for the proposed Coulomb control system are calculated analogous to traditional propulsion systems and it is shown that required forces can be created with milliwatts of power, can be controlled on a millisecond time scale, and imply specific impulse that can be as high as  $10^{13}$  seconds.

### 1. Introduction

Swarms of microsatellites are envisioned as an alternative to traditional large spacecraft. Such swarms, acting collectively as virtual satellites, will benefit from the use of cluster orbits where the satellites fly in a close formation.<sup>1</sup> The formation concept, first explored in the 1980's to allow multiple geostationary satellites to share a common orbital slot,<sup>2,3</sup> has recently entered the era of application with many missions slated for flight in the near future. The promised payoff of formation-flying has recently inspired a large amount of research in an attempt to overcome the rich technical problems. A variety of papers can be found in the proceedings of the 1999 AAS/AIAA Space Flight Mechanics Meeting,<sup>4,5,6</sup> the 1998 Joint Air Force/MIT Workshop on Satellite Formation Flying and Micro-Propulsion,<sup>7</sup> a recent textbook on micropropulsion,<sup>8</sup> and numerous other sources.<sup>9,10,11,12,13,14,15</sup>

Relative position control of multiple spacecraft is an enabling technology for missions seeking to exploit satellite formations. Of the many technologies that must be brought to maturity in order to realize routine formation flying, perhaps the most crucial is the spacecraft propulsion system. The most challenging propulsion system demands will be made in close formations where inter-vehicle spacing may be as small as five or ten meters. Apart from the obvious danger of collision, exhaust plume contamination of sensitive instruments is a legitimate concern for spacecraft in a formation. In close proximity, the propellant emitted by such devices as micro-PPT's (vaporized Teflon), FEED (ionized cesium), or colloid thrusters (liquid glycerol droplets doped with NaI) may impinge upon neighboring vehicles with the potential to damage payloads. To worsen the problem, orbital mechanics for many clusters of interest mandate continuous thruster firings pointed directly towards other vehicles in the formation.

Work reported in this paper represents what the authors believe is a new mode of propulsive interaction between spacecraft in close formation.

The interaction arises from Coulomb forces between vehicles in a swarm of electrically charged spacecraft. Forces can be produced as a result of natural charging due to space plasma interaction. In the case of natural charging, the Coulomb forces represent perturbations on vehicles in the formation, which on-board thrusters must counteract. It will be shown in following sections that, in high orbits such as GEO, natural charging is capable of producing forces and torques with magnitudes comparable to those of microthrusters over separation length scales of tens of meters.

In addition to an analysis of parasitic Coulomb forces from space plasma interaction, this paper reports on an exploratory study to examine the feasibility of using Coulomb forces to maintain rigid satellite formations in high earth orbit. In the Coulomb control scenario, on-board power would be used to actively charge vehicles in a swarm to affect formation-keeping forces and possibly act as an emergency collision avoidance system. The use of Coulomb forces would provide nearly propellantless formation control with little or no risk of exhaust plume contamination of neighboring vehicles. Analytic and numeric calculations show that the dynamic equations permit rigid Coulomb formations in orbit and that these formations are controllable in three dimensions. It is also shown that the amount of on-board power required to affect charge control is negligible and that continuous thrust variation is possible over rapid time scales.

## 2. Coulomb Perturbations

### 2.1. GEO Plasma Environment

The Coulomb force between two point charges in the presence of a plasma is represented by the typical inverse square relationship modified by an exponential term to account for the Debye shielding

$$\text{Eqn. 1} \quad F_{1,2} = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_2}{d_{1,2}^2} \exp\left(\frac{-d_{1,2}}{I_d}\right),$$

Where  $q_1, q_2$  are the charges,  $d_{1,2}$  is the separation distance, and  $I_d$  is the plasma Debye length. The potential for interaction between charged spacecraft is then limited to regimes in which the separation between vehicles is less than the plasma Debye length. Since the Debye length in low-Earth orbit (LEO) is on the order of centimeters,

Coulomb interaction will be present only in high-earth orbits. Analysis in this study will be limited to geostationary Earth orbit (GEO) to take advantage of the wealth of plasmaphysical data available for GEO conditions.

An excellent discussion of the space plasma environment can be found in the textbook by Hastings and Garret,<sup>16</sup> some of which is repeated here for convenience. A spacecraft at GEO is at the edge of plasmopause. This collisionless plasma does not follow a single Maxwellian distribution. Instead, plasma parameters must be measured experimentally. The particle detectors on the ATS<sup>17,18,19</sup> and SCATHA<sup>20</sup> spacecraft have measured plasma variations between 5-10 eV and 50-80 eV approximately, for 50 complete days at 1 to 10 minute resolution from 1969 through 1980, bracketing one solar cycle.

Garrett and Deforest<sup>17</sup> fitted an analytical two-temperature model to data collected over 10 different days from ATS-5 spacecraft between 1969 and 1972. These data were selected in such a way to show a wide range of geomagnetic activity including plasma injection events (i.e. sudden appearance of dense, relatively high energy plasma at GEO occurring at local midnight). The model gives reasonable and consistent representation of the variations following a substorm injection event at GEO. The parameters for this model during average GEO conditions are shown in Table 1 with "Worst-case" GEO conditions given in Table 2. Calculations based on these parameters show that the Debye length at GEO ranges from about 140 m to greater than 1400 m.

Parameter	Electrons	Ions
$n_1$ ( $m^{-3}$ )	$0.78 \times 10^6$	$0.19 \times 10^6$
$kT_1/e$ (eV)	550	800
$n_2$ ( $m^{-3}$ )	$0.31 \times 10^6$	$0.39 \times 10^6$
$kT_2/e$ (eV)	$8.68 \times 10^3$	$15.8 \times 10^3$

Table 1. Average GEO environment.<sup>17</sup>

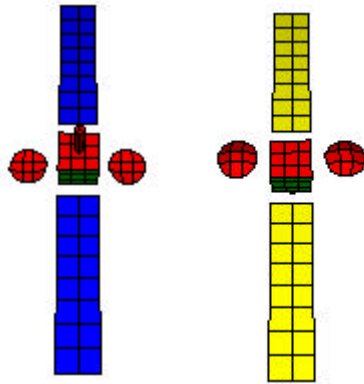
Parameter	Electrons	Ions
$n_1$ ( $m^{-3}$ )	$1 \times 10^6$	$1.1 \times 10^6$
$kT_1/e$ (eV)	600	400
$n_2$ ( $m^{-3}$ )	$1.4 \times 10^6$	$1.7 \times 10^6$
$kT_2/e$ (eV)	$2.51 \times 10^4$	$2.47 \times 10^4$

Table 2. Worst-case GEO environment.<sup>17</sup>

### 2.2. Modelling Spacecraft Charging

Spacecraft charging, especially differential charging, has been of prime concern to spacecraft designers because of its detrimental effects such as

electrostatic discharge in spacecraft and spacecraft subsystems. The NASA Space Environment Effects (SEE) Interactive Spacecraft Charging Handbook<sup>21</sup> is one of the tools available to model the plasma environment and spacecraft charging. The SEE tool allows specification of plasma parameters, spacecraft size, materials, and charging time, whereupon the program predicts potentials of elements on the spacecraft surface. The transient response of a spacecraft in plasma is calculated by modeling the spacecraft – ambient plasma system as an equivalent electric circuit. The equilibrium charge of a vehicle is obtained from the SEE model as the transient solution reaches steady state.



Component	Color	Size	Material
Chassis	Red Green	1mx1mx1m	Kapton OSR
Solar Arrays	Blue Yellow	1mx4m	Solar Cells Black Kapton
Antennae	Red	1-m-dia.	Kapton
Omni Antenna	Red	0.2-m-dia. 1-m-long	Kapton

Figure 1. Spacecraft model used in SEE code seen from sun direction (left) and from opposite direction (right) with representative materials of each component.

The SEE code was used to predict spacecraft charging, with the intent of using the results of the SEE predictions to calculate Coulomb perturbations. The default spacecraft materials of the SEE code were used for these tests, which are shown in Figure 1. The SEE code has three pre-loaded GEO plasma environments, namely Worst-Case environment, ATS-6 environment and 4Sept97 environment. The specifications of these environments are given in Table 3. A representative solution from the SEE tool is shown in Figure 2. This figure shows the surface potential distribution (in units of volts) as a color-contour map.

Parameter	Worst-case	ATS-6	4-Sept.-97
$n_e$ ( $m^{-3}$ )	$1.2 \times 10^6$	$1.22 \times 10^6$	$3.00 \times 10^5$
$T_e$ (eV)	$1.2 \times 10^3$	$1.6 \times 10^6$	$0.4 \times 10^3$
$n_i$ ( $m^{-3}$ )	$2.36 \times 10^7$	$2.36 \times 10^7$	$0.30 \times 10^7$
$T_i$ (eV)	$2.95 \times 10^3$	$2.95 \times 10^3$	$0.40 \times 10^3$

Table 3. Parameters of the three pre-loaded GEO plasma environments in the SEE Interactive Spacecraft Charging Handbook.

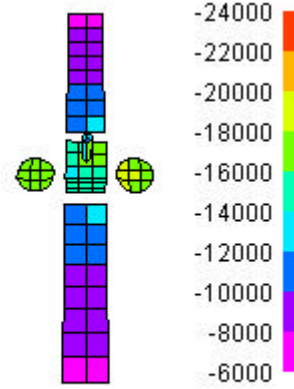


Figure 2. SEE prediction of spacecraft surface potentials for vehicle subject to worst-case GEO plasma environment in non-eclipse conditions. Units are volts.

### 2.3. Natural Forces and Torques

The output of the SEE model includes a value of potential for each finite element of the spacecraft surface. Using a vacuum Green's function

$$\text{Eqn. 2} \quad 4\pi\epsilon_0 V_i = \sum_j \frac{q_j}{\bar{r}_{ij}}$$

it is possible to calculate the equivalent point charge at the center of each finite element that is self consistent with the potential distribution. The interaction between two charged spacecraft can then be modeled as a superposition of the interactions of the equivalent point charges. For example, the collection of charges on spacecraft 'B' will produce an electric field at a point 'i' on spacecraft 'A' according to

$$\text{Eqn. 3} \quad \vec{E}_i = \frac{1}{4\pi\epsilon_0} \sum_{j=1}^{nB} \frac{q_j \bar{r}_{ij} \exp\left(-\frac{|\bar{r}_{ij}|}{I_d}\right)}{|\bar{r}_{ij}|^3}$$

where  $n_B$  is the number of elements (charges) on spacecraft 'B'. If there is an electric charge  $q_i$  at 'i' it will experience a force  $\vec{F}_i = q_i \vec{E}_i$ . The net force on spacecraft 'A' resulting from the  $n_B$  charges on spacecraft 'B' is then simply

$$\text{Eqn. 4} \quad \vec{F}_{AB} = \sum_{i=1}^{n_B} \vec{F}_i.$$

Using the same equivalent point charge method it is also straightforward to compute the moment on the charged spacecraft 'A' about its geometric center resulting from the distribution of charges on spacecraft 'B'.

The Coulomb interaction between closely spaced satellites was performed using the SEE code to predict the charging of representative spacecraft. Two identical vehicles, with material properties and physical dimensions as shown in Figure 1, were assumed to be in GEO separated along the direction of orbital velocity to mimic a leader-follower type of formation. The relative orientation of the vehicles is indicated in Figure 3.

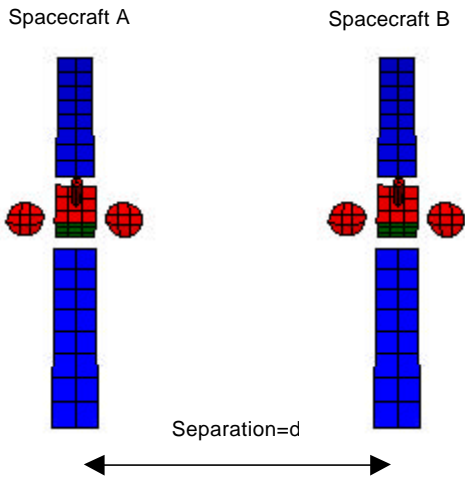


Figure 3. Relative orientation of two-spacecraft leader-follower formation for natural Coulomb interaction study. The vehicles are displaced only along the orbital velocity direction.

Three GEO plasma environments (see Table 3) were used to predict the spacecraft charging in both eclipse and non-eclipse conditions due to natural plasma interaction. The forces and torques were computed based on the SEE results for a range of spacecraft separations from 10 to 100 meters. The resulting inter-spacecraft forces and torques about geometric center are shown in Figure 4. The Coulomb interaction forces were found to be as large as 1 mN for spacecraft 10 m apart in the

ATS-6 eclipse environment, with all environments except the 4-Sept.-97 case showing interaction forces greater than 10  $\mu$ N at the closest spacing. The decay in force with separation is not purely  $1/r^2$  due to the finite size effects of the vehicles. At the largest spacing considered (100 m) the inter-spacecraft forces vary from  $10^{-10}$  N up to about 100 nN, depending upon the orbital conditions used in the SEE prediction. The electric-dipole-induced torques were found to be as large as 100  $\mu$ N-m for the closest spacing in the ATS-6 Eclipse conditions, falling as low as  $10^{-10}$  N-m for the 4-Sept.-97 case at 100-m spacing.

A surprising result of the Coulomb interaction study was that the magnitude of the inter-spacecraft forces is comparable with and may exceed that of candidate micropropulsion systems proposed for formation keeping. One method for eliminating the parasitic Coulomb forces and torques would be to actively control the spacecraft charge and ensure that vehicle potential remains very near the ambient plasma potential. This could be performed using a hollow-cathode plasma contactor as used on the International Space Station (ISS). Another interesting concept, however, is to use an ion or electron emitter to actively control the spacecraft charge and, thus, the Coulomb force to affect relative propulsion within a formation. The feasibility of forming Coulomb-stabilized formations will be the focus of Section 3.

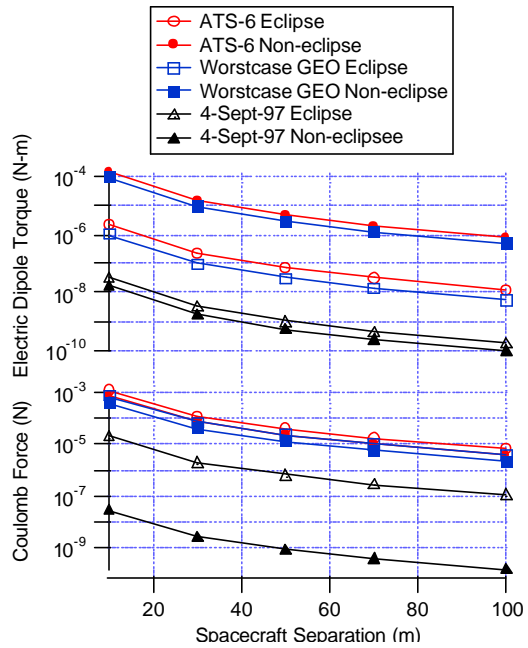


Figure 4. Predicted inter-spacecraft Coulomb force for two identical satellites in a GEO leader-follower formation as a function of the formation spacing.

### 3. Formation Dynamics

This section presents a summary of Coulomb formation dynamics that can be found in much more detail in the references.<sup>22</sup> Analytic methods are presented for determining static (rigid) formations of spacecraft flying in close proximity using inter-vehicle Coulomb forces to offset orbital perturbations. The vehicles are approximated as spheres and it is assumed that the plasma Debye length is much longer than a typical inter-vehicle separation. Such conditions are a reasonable approximation of formations in GEO with vehicle spacing on the order of tens of meters.

#### 3.1. Dynamic Equations

As in many formation studies, Hill's equations are used here to describe the motion of spacecraft in a formation relative to a reference point that is assumed to maintain a circular Keplerian orbit. With an application slanted towards separated spacecraft interferometry, the central reference vehicle is referred to as a "combiner" where the surrounding vehicles are called "collectors". It is assumed that the combiner has its own station keeping system, but the collectors do not. Thus the only external forces on the collectors are the Coulomb interactions between them and the combiner.

Within the Hill's system, the motion of the  $i$ 'th collector with respect to the central combiner-fixed coordinate system can be written for a formation of  $n$  vehicles interacting via Coulomb forces as

Eqn. 5

$$\begin{aligned}\ddot{x}_i - 2\Omega\dot{y}_i - 3\Omega^2x_i &= \frac{k_c}{m_i} \sum_{j=0}^n \frac{(x_i - x_j)}{|\bar{p}_i - \bar{p}_j|^3} q_i q_j \\ \ddot{y}_i + 2\Omega\dot{x}_i &= \frac{k_c}{m_i} \sum_{j=0}^n \frac{(y_i - y_j)}{|\bar{p}_i - \bar{p}_j|^3} q_i q_j \\ \ddot{z}_i + \Omega^2z_i &= \frac{k_c}{m_i} \sum_{j=0}^n \frac{(z_i - z_j)}{|\bar{p}_i - \bar{p}_j|^3} q_i q_j \\ i &= 1, \dots, n \text{ and } i \neq j\end{aligned}$$

where  $\bar{p}_i$  denotes the position vector of the  $i$ 'th vehicle,  $m_i$  is the spacecraft mass,  $\Omega$  is the orbital angular velocity,  $q_i$  is the vehicle charge in Coulombs, and  $k_c=1/4\pi\epsilon_0$  is Coulomb's constant.

The coordinate notation is such that the  $y$  direction is along the orbital velocity vector,  $x$  is in the zenith-nadir direction, and  $z$  is normal to the orbit plane. The axis system is shown in Figure 5. The interesting difference between this set of equations and the typical Hill's system used in formation studies lies in their coupling: the dynamics of any vehicle in the formation is influenced simultaneously by all of the other vehicles in the group.

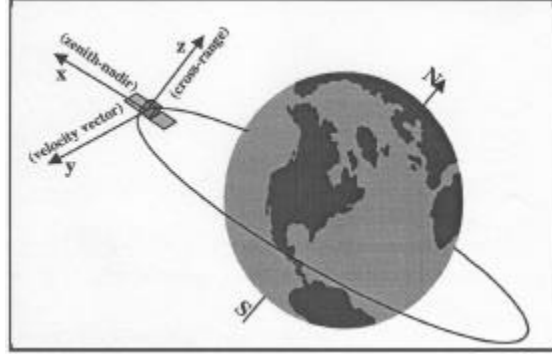


Figure 5. Illustration of combiner-fixed relative coordinate system used in the Hill's equation formulation (Figure reprinted from Ref. 23).

A static (rigid) formation geometry must satisfy the Hill's system of Eqn. 5 with zero relative velocity and acceleration for each vehicle ( $\dot{x} = \ddot{x} = \dot{y} = \ddot{y} = \dot{z} = \ddot{z} = 0$ ). The nature of the Coulomb coupling of the system stipulates that control forces can only be applied along the lines "connecting" spacecraft coordinates. Thus, the goal is to find suitable formation geometries such that the vector sum of the Coulomb forces is sufficient to enable solution of the Hill's system with velocities and accelerations set to zero. The equilibrium system of equations is then found to be

$$-3\Omega^2 x_i = \frac{k_c}{m_i} \sum_{j=0}^n \frac{(x_i - x_j)}{|\bar{p}_i - \bar{p}_j|^3} q_i q_j$$

$$0 = \frac{k_c}{m_i} \sum_{j=0}^n \frac{(y_i - y_j)}{|\bar{p}_i - \bar{p}_j|^3} q_i q_j$$

Eqn. 6

$$\Omega^2 z_i = \frac{k_c}{m_i} \sum_{j=0}^n \frac{(z_i - z_j)}{|\bar{p}_i - \bar{p}_j|^3} q_i q_j$$

$$i = 1, \dots, n \text{ and } i \neq j$$

### 3.2. Formation Geometries

The method chosen for exploring the behavior of Eqn. 6 was to assume geometric formations and search for solution sets. Although numerous formations were studied in the reference work,<sup>22</sup> discussion here will be limited to three canonical cases. The first case considered was a simple linear three-vehicle formation with vehicles separated by  $L$ . With the combiner fixed to the origin, three sub-formations follow depending upon orientation with the axis system. The linear formations are shown in Figure 6.

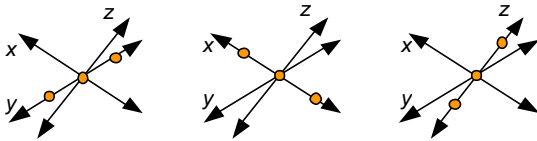


Figure 6. Linear three-spacecraft formations.

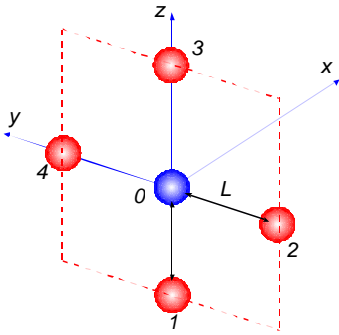


Figure 7. Illustration of the five-satellite formation.

The second formation studied was a simple two-dimensional formation that suggests an application for Earth observation through separated spacecraft interferometry. This configuration is shown in Figure 7, where the four collectors are distributed evenly about the Hill's  $y$  and  $z$  axes to create a five-spacecraft planar formation. The third formation represented a three-dimensional configuration, with six collectors distributed evenly about the principal axes with spacing  $L$ . The three-dimensional formation is shown in Figure 8.

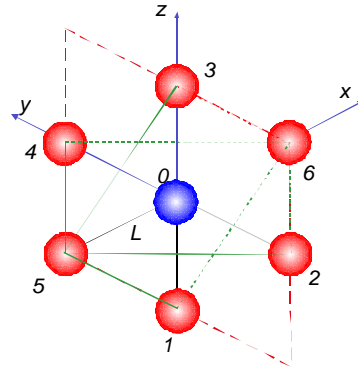


Figure 8. Illustration of seven-satellite formation.

### 3.3. Equilibrium Formation Solutions

Although results will be presented for all of the formations outlined above, the analytic solution method will be presented only for the representative three-spacecraft  $x$ -axis aligned case, which we call the "Coulomb tether" due to its similarity with physical tether configurations. For three spacecraft aligned along the combiner coordinate frame's  $x$ -axis as shown in Figure 6,  $n=3$  and the following relative displacement constraints hold

$$\begin{aligned} x_1 &= L \\ x_2 &= -L \\ y_1 &= y_2 = z_1 = z_2 = 0 \end{aligned}$$

Eqn. 7

where  $L$  is the distance from the combiner to either collector. Forming all six of the equilibrium equations from Eqn. 6 and eliminating duplicate equations leaves only two conditions:

$$\text{Eqn. 8} \quad \frac{k_c q_1 q_2}{4L^2} + \frac{k_c q_1 q_0}{L^2} + 3m_1 L \Omega^2 = 0$$

$$\text{Eqn. 9} \quad \frac{k_c q_1 q_2}{4L^2} + \frac{k_c q_2 q_0}{L^2} + 3m_2 L \Omega^2 = 0$$

If we further assume that the collectors have equal mass,  $m=m_1=m_2$  and introducing the normalized charges defined by

$$\text{Eqn. 10} \quad q_{in} \equiv \frac{q_i}{\sqrt{m_i L^3}}$$

allows Eqn. 8 and Eqn. 9 to be written without explicit mass and length dependencies

$$\text{Eqn. 11} \quad \begin{aligned} \frac{k_c q_{1n} q_{2n}}{4} + k_c q_{1n} q_{0n} + 3\Omega^2 &= 0 \\ \frac{k_c q_{1n} q_{2n}}{4} + k_c q_{2n} q_{0n} + 3\Omega^2 &= 0 \end{aligned}$$

where the subscript  $n$  denotes a normalized quantity. Assuming spherical spacecraft we can use Gauss' Law to relate the vehicle surface potential in volts to the equivalent encircled point charge  $q_i$  as  $V_i = q_i / 4\pi\epsilon_0 r_i$  where  $r_i$  is the spacecraft radius. In a loosely analogous fashion we will define a normalized voltage  $V_{in}$  based on the normalized charge  $q_{in}$

$$\text{Eqn. 12} \quad V_{in} \equiv k_c q_{in}$$

then the equilibrium equations of Eqn. 11 are

$$\text{Eqn. 13} \quad \begin{aligned} V_{1n} V_{2n} + 4V_{1n} V_{0n} + 12k_c \Omega^2 &= 0 \\ V_{1n} V_{2n} + 4V_{2n} V_{0n} + 12k_c \Omega^2 &= 0 \end{aligned}$$

These are readily solved analytically. Given a suitable combiner spacecraft normalized voltage,  $V_{0n}$ , the two collector normalized voltages must be equal and are

$$\text{Eqn. 14} \quad V_{1n} = V_{2n} = -2V_{0n} \pm 2\sqrt{V_{0n}^2 - 3k_c \Omega^2}$$

where the combiner normalized voltage must satisfy the constraint

$$\text{Eqn. 15} \quad V_{0n}^2 - 3k_c \Omega^2 \geq 0$$

Knowing the actual collector mass,  $m$ , radius,  $r$ , separation,  $L$ , and the orbital angular rate  $\Omega$ , the equilibrium collector physical voltage can be obtained from Eqn. 14 and the normalization relationship

$$\text{Eqn. 16} \quad V_{in} = \frac{V_i r_i}{\sqrt{m L^3}}$$

where the quantity  $V_i r_i$  (which is equal to  $q_i / 4\pi\epsilon_0$  by Gauss' Law) is called the *reduced charge* of spacecraft  $i$ . The normalized collector voltages, obtained from Eqn. 14, are shown in Figure 9 as a function of the combiner voltage,  $V_{0n}$ . The angular rate  $\Omega$  is for a geosynchronous orbit,  $\Omega = 7.2915 \times 10^{-5}$  rad/s. Similar analytic methods were used to compute the y- and z-axis aligned formations. These solutions are also shown in Figure 9.

The three-spacecraft solution set yield some interesting insight. The most trivial case is the y-axis geometry. For this alignment within the Hill's system, a solution set is possible where all of the vehicles are uncharged. Indeed, if any vehicle has a non-zero charge, the net effect of this charge such that the net force is always zero. The z-axis formation permits solutions where the central combiner has no charge, however there is no solution where all vehicles are neutral and Coulomb forces are required to maintain the static equilibrium. Lastly, the x-axis formation does not permit any solutions with uncharged spacecraft. There is a clear minimum magnitude for  $V_{0n}$  (about 20 in the normalized units) below which no solution is possible.

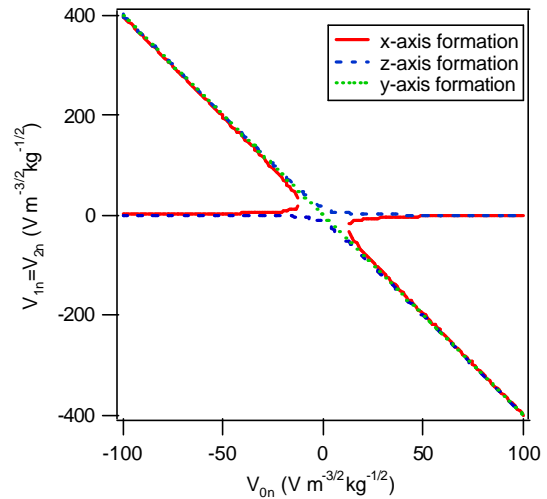


Figure 9. Analytic solution set for equilibrium three-spacecraft linear formations.

The analytic method was extended to the five-spacecraft two-dimensional formation and the seven-spacecraft three-dimensional formation.

When the constraints of the two- and three-dimensional formations were applied to the equilibrium requirements of Eqn. 6, the result was, respectively, eight and eighteen unique necessary conditions for equilibrium.

The two-dimensional five-spacecraft formation yielded two families of solutions. The first family was somewhat trivial and produced solutions where two of the vehicles had zero charge, such that the remaining vehicles assumed the same three-spacecraft linear solutions described above. The second family of solutions was found by forcing the symmetry condition that  $q_1=q_3$  and  $q_2=q_4$ ; this reduced the set of eight unique conditions to two.

Eqn. 17

$$4V_{0n}V_{3n} + V_{3n}^2 + 2\sqrt{2}V_{3n}V_{4n} - k_c\Omega^2 = 0$$

$$4V_{0n}V_{4n} + V_{4n}^2 + 2\sqrt{2}V_{3n}V_{4n} = 0$$

Eqn. 17 can be solved conditionally such that  $q_1$  (and, hence  $q_3$ ) have the form  $q_1=q_1(q_4)$  and  $q_0=q_0(q_1, q_4)$ . The resulting solution set is shown in Figure 10. The solid and dashed lines represent consistent solutions within the set. For example, if  $V_{2n}=V_{4n}=-50$ , then *either*  $V_{1n}=V_{3n}=-10$  with  $V_{0n}=20$ , or  $V_{1n}=V_{3n}=-50$  with  $V_{0n}=50$

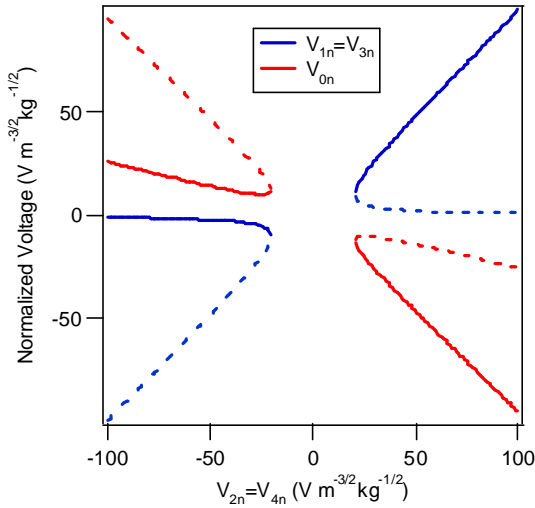


Figure 10. Solution set for equilibrium five-spacecraft two-dimensional formation.

The seven-spacecraft equilibrium solution was found by requiring the symmetry condition that  $q_1=q_3$ ,  $q_2=q_4$ , and  $q_5=q_6$ , reducing the set of eighteen conditions to three unique equations:

Eqn. 18

$$\frac{k_c}{L^2}q_0q_1 + \frac{\sqrt{2}}{2L^2}k_cq_1q_2 + \frac{k_cq_1^2}{4L^2} + \frac{\sqrt{2}}{2L^2}k_cq_1q_5 - mL\Omega^2 = 0$$

$$\frac{1}{L^2}k_cq_0q_2 + \frac{\sqrt{2}}{2L^2}k_cq_1q_2 + \frac{1}{4L^2}k_cq_2^2 + \frac{\sqrt{2}}{2L^2}k_cq_2q_5 = 0$$

$$\frac{k_cq_0q_5}{L^2} + \frac{\sqrt{2}}{2L^2}k_cq_2q_5 + \frac{\sqrt{2}}{2L^2}k_cq_1q_5 + \frac{k_cq_5^2}{4L^2} + 3mL\Omega^2 = 0$$

Assuming a value of  $q_2$  (or similarly  $q_4$ ), Eqn. 18 can be solved analytically such that the collectors' charges are functions of  $q_2$

Eqn. 19

$$q_1 = q_3 = q_1(q_2)$$

$$q_5 = q_6 = q_5(q_2)$$

The combiner charge,  $q_0$ , can then be expressed as  $q_0 = q_0(q_1, q_2, q_5)$ . A plot of the equilibrium solution set is shown in Figure 11. In the figure, solid and dashed lines represent consistent solutions within the set. For example, if  $V_{2n}$  is  $-50$ , then the solution is *either*  $V_{1n}=V_{3n}=-50$ ,  $V_{5n}=V_{6n}=-60$ , with  $V_{0n}=75$  or  $V_{1n}=V_{3n}=-10$ ,  $V_{5n}=V_{6n}=8$ , with  $V_{0n}=10$ .

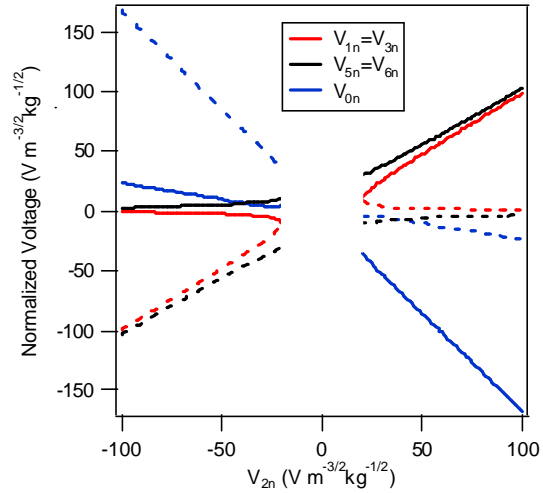


Figure 11. Equilibrium solution set for seven-spacecraft three-dimensional formation.

### 3.4. Physical Estimates

The normalized and reduced variables utilized in the analytic solution method can be used to estimate physically meaningful parameters. It is apparent from a study of Figure 9, Figure 10, and Figure 11 that an infinite number of possible vehicle charge states (voltages) exist for any of the



formations. One method of selecting an “optimal” set of parameters was based on minimizing the overall charging required by the formation. Using the sum of the squares of the vehicle charges as a cost function, the optimum solution for each formation was calculated analytically. The optimal solutions can be used to provide estimates of physical parameters. For discussion sake, it is assumed here that the vehicles each have identical mass of 150 kg, and that the inter-spacecraft separation for all of the formations was  $L=10$  m. From the optimum solutions the value of reduced charge,  $V_i r_i$ , can be calculated. The reduced charge can be thought of as the surface potential required for a spherical vehicle with 1-m-radius, with scaling to different radii vehicles a straightforward calculation. The magnitude of optimal vehicle reduced charge for each formation is presented in Table 4.

	$V_0 r_0$	$V_1 r_1$	$V_2 r_2$	$V_3 r_3$	$V_4 r_4$	$V_5 r_5$	$V_6 r_6$
<b>3-x</b>	13.8	13.8	13.8				
<b>3-z</b>	2.39	2.39	2.39				
<b>5</b>	4.78	3.96	7.92	3.96	7.92		
<b>7</b>	12.7	3.96	7.92	3.96	7.92	11.2	11.2

Table 4. Minimum total charge equilibrium solutions for each of the formations. For the linear formations, only the x-axis aligned and z-axis aligned three-spacecraft solutions are shown. The units of the vehicle reduced charges ( $V_i r_i$ ) are in kV-m. The numbers can be thought of as the surface potential of a 1-m-radius spherical spacecraft in kilovolts.

#### 4. Propulsion Figures of Merit

The purpose of this section is to evaluate some fundamental performance metrics of a Coulomb control system on a spacecraft formation. Aspects such as control force, input power, required consumable mass, and environment interaction will be calculated for a simple two-spacecraft system.

##### 4.1. Required Power

An isolated spacecraft will assume an equilibrium potential (voltage) such that the net environmental current due to plasma and photoelectron emission is zero. It is possible to change the vehicle potential by emitting charge from the spacecraft. For example, if it is desired to drive the spacecraft potential lower than equilibrium (more negative), the emission of positive ions from the vehicle will cause a net surplus of on-board electrons and a lowering of the potential. In order to emit such a current, the charges must be ejected from the vehicle with

sufficient kinetic energy to escape the spacecraft potential well. Thus, if the vehicle is at a (negative) potential  $-V_{SC}$ , then ions must be emitted from a source operating at a power supply voltage,  $V_{PS}$ , greater than  $|V_{SC}|$ .

Basic concepts can be used to calculate the power required to maintain the spacecraft at some steady state potential. To maintain the spacecraft at a voltage of  $|V_{SC}|$ , current must be emitted in the amount of  $|I_e| = 4\pi r^2 |J_p|$ , where  $J_p$  is the current density to the vehicle from the plasma, using a power supply having voltage of at least  $|V_{PS}| = |V_{SC}|$ . Quantitatively,

$$\text{Eqn. 20} \quad P = |V_{SC} I_e|.$$

For a two-spacecraft formation with each vehicle using Power  $P$ , the total system power is just the sum of the individual power to each vehicle. Assuming spherical spacecraft and using Gauss' Law to relate the surface potential to the encircled point charge, it is possible to relate the Coulomb force (thrust) on a vehicle to the emission current and the required power

$$\text{Eqn. 21} \quad F_C = 4\pi e_0 e^{-d/l_d} \frac{r_A r_B P^2}{d^2 I_{eA} I_{eB}}$$

where  $r_{A(B)}$  and  $I_{eA(B)}$  are the radius and emission current of spacecraft A(B), and  $d$  is the vehicle separation.

Eqn. 21 yields the power required to produce a steady-state thrust for a given emission current. Since the space environment will be constantly changing (and, hence the emission current to maintain steady state), it is important to calculate the required power to affect a change in potential from some initial value to a desired steady state value. In the pedagogical analysis here, the capacitance of the spherical spacecraft can be used to estimate the power required for a change in voltage (thrust). Using an equivalent circuit model where  $dV_{SC}/dt=I/C$ , the rate of change of spacecraft potential can be related to the current absorbed from the plasma and the emitted control current

$$\text{Eqn. 22} \quad \frac{dV_{SC}}{dt} = \frac{4\pi r^2 J_p + I_e}{4\pi e_0 r}$$

where  $J_p$  is the absorbed plasma current density. The value of  $J_p$  can easily be evaluated using traditional plasma probe theory for a sphere and will take the form  $J_p = J_p(V_{SC}, T_e, T_i)$ . If  $V_{final}$  is the desired steady state spacecraft voltage, then the emission current power supply must have power  $P = I_e V_{final}$ . Substituting  $I_e = P/V_{final}$  into Eqn. 22 and

using an analytic form for the plasma current, an explicit equation is obtained of the form

$$\text{Eqn. 23} \quad \frac{dV_{SC}}{dt} = f(P, V_{SC}, T_e, T_i, r)$$

which can be numerically integrated to produce a function

$$\text{Eqn. 24} \quad V_{SC} = V_{SC}(T_e, T_i, r, P, t).$$

As a numerical example, Figure 12 shows a plot of the function obtained for Eqn. 24 assuming a 1-m-diameter spacecraft charging from  $V_{SC}=0$  to  $V_{final}=6$  kV in the average GEO plasma environment. From this plot it is evident that only 200 mW of power is required to change the spacecraft potential by 6 kV within 8 msec.

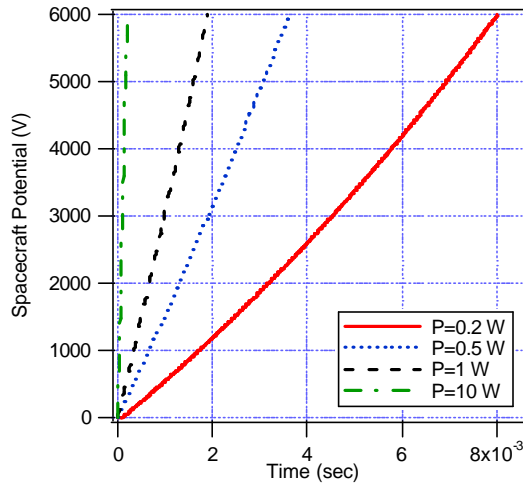


Figure 12. Numerical integration of the transient response of a 1-m-diameter model spacecraft in an average GEO plasma as a function of the power in the emitted control beam.

#### 4.2. Mass Flow Rate and $I_{sp}$

Mass flow rate is defined by the rate of gaseous ions expelled out per unit time to maintain potential of the vehicle. As electrons have negligible mass we can say that mass flow rate of electrons is negligible and thus driving the potential positive requires zero mass flow. If  $I_e$  is the emission current constituting ions,  $m_{ion}$  is the mass of ion, and  $q_{ion}$  is the charge, then mass flow rate is given by

$$\text{Eqn. 25} \quad \dot{m} = \frac{I_e m_{ion}}{q_{ion}}.$$

Since the only purpose of the ion emission is to carry charge (and not momentum) from the vehicle, it makes sense to use the least massive ions that are practical. For a two spacecraft combination, propellant mass flow rate will be the sum of mass flow rates for individual spacecraft and can be related to their individual emission currents

$$\text{Eqn. 26} \quad \dot{m}_{Total} = \frac{m_{ion}}{q_{ion}} (I_{ea} + I_{eb}).$$

A common performance parameter used for propulsion systems is specific impulse  $I_{sp}$ . This parameter compares the thrust derived from a system to the required propellant mass flow rate.<sup>24</sup> Although  $I_{sp}$  is traditionally used as a parameter to evaluate momentum transfer (rocket) systems, we can use the formal definition to compare the Coulomb system. For a Coulomb control system the specific impulse  $I_{sp}$  is given by

$$\text{Eqn. 27} \quad I_{sp} = \frac{F}{\dot{m}_{Total} g_0}$$

Since Coulomb force calculations are meaningless for a single vehicle, we will treat the system as two separate vehicles, each subject to a force of  $F_c$  given by Eqn. 21, so that the sum of the forces experienced by all spacecraft in the formation is  $F=2F_c$ .

Eqn. 28

$$I_{sp} = \frac{8pe_0 e^{-d/1d} q_{ion}}{g_0 m_{ion}} \frac{r_A r_B P^2}{d^2 I_{eA} I_{eB} (I_{eA} + I_{eB})}$$

Where  $g_0$  is the gravitational constant. If  $r_A = r_B = r_{sc}$ , and  $I_e = I_{eA} = I_{eB}$ , then Eqn. 28 becomes,

$$\text{Eqn. 29} \quad I_{sp} = \frac{4pe_0 e^{-d/1d} q_{ion} r_{SC}^2 P^2}{g_0 m_{ion} d^2 I_e^3}.$$

Note that, unlike a rocket system, the definition of  $I_{sp}$  of a coulomb system is meaningless for a single vehicle. For a formation of two spacecraft, Eqn. 29 indicates that the specific impulse of the formation is a function of the radii of the spacecraft, power supplied to the ion (electron) gun, the separation between the two spacecraft, the emission currents of both vehicles, and the mass of the charge carriers,  $m_{ion}$ .

Consider a two-spacecraft formation with identical 1-m-diameter vehicles in the average GEO plasma environment charged to the same

negative potential. In order to reach and maintain this negative potential, the vehicles must emit an ion current. Consequently, the spacecraft will attract ion saturation current from the plasma, so  $I_e$  must be equal to the plasma ion saturation current for steady state. It is apparent that light ions will provide the most efficient  $I_{sp}$ , so assume that the emitted species is  $H^+$ . Calculated values of specific impulse for each vehicle in the formation is shown in Figure 13 for various system input power levels. For 1 mW systems with vehicle separation on the order of 20 m,  $I_{sp}$  values of  $10^4$  seconds are obtained, with values increasing to  $10^{10}$  sec for just 1 W of power. It should be noted that for a positive vehicle potential the emitted species would be electrons and, thus, the calculated values of  $I_{sp}$  would be a factor of 2,000 greater.

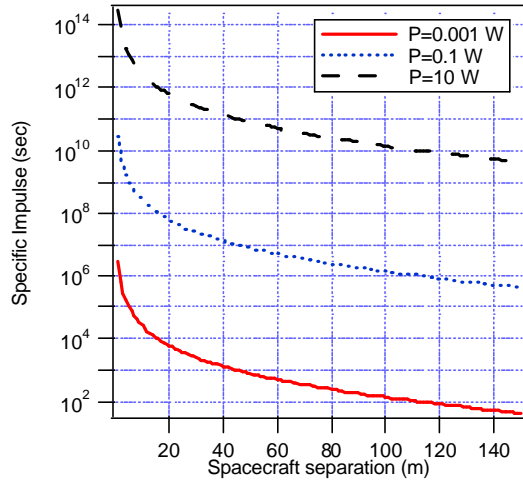


Figure 13. Specific impulse for a two-spacecraft Coulomb formation as a function of spacecraft separation,  $d$ , and input power,  $P$ .

### 4.3. Emission Current Jet Force

Generating net charge on a spacecraft for Coulomb force requires the emission of current. In principle, the charge is carried away from the vehicle by particles with non-zero mass. Such mass ejection results in a momentum jet force on the vehicle as in a traditional electric propulsion thruster. In the case of electron emission, the mass of the charge carriers is insignificant and the resulting jet force is negligible. Ion emission, however, may produce a significant reaction force. It is instructive to consider how the Coulomb force between spacecraft compares with the momentum

reaction on the vehicle induced by the beam of ion current.

The reactive thrust force of an ejected mass flow is computed as

$$\text{Eqn. 30} \quad F_J = \dot{m}u_e,$$

where  $\dot{m}$  is the ejected mass flow rate and  $u_e$  is the exhaust velocity at which the mass is emitted. Assuming steady state Coulomb force generation, the ions will be electrostatically accelerated through a spacecraft potential of  $V_{SC}$ , such that

$$\text{Eqn. 31} \quad u_e = \sqrt{\frac{2q_{ion}V_{SC}}{m_{ion}}}.$$

With this simplification and recognizing that the mass flow is related to the emission current, the momentum jet force of the emitted ion current is

$$\text{Eqn. 32} \quad F_J = I_e \sqrt{\frac{2m_{ion}V_{SC}}{q_{ion}}}.$$

The jet force can also be written in terms of the input power to the emission system as

$$\text{Eqn. 33} \quad F_J = \sqrt{\frac{2m_{ion}PI_e}{q_{ion}}}.$$

We can compare the magnitude of the jet reaction force with the induced Coulomb force between two vehicles. Assume identical spacecraft charged to the same value of  $V_{SC}$ . From Eqn. 21 and Eqn. 33 we can write the ratio of  $F_C/F_J$  (taking  $F_C$  as the total Coulomb force on both vehicles) in terms of the input power as

$$\text{Eqn. 34} \quad \frac{F_C}{F_J} = 4\sqrt{2}pe_0 \sqrt{\frac{q_{ion}}{m_{ion}}} \frac{r_A r_B P^{3/2} e^{-d/l_d}}{I_{eA} I_{eB} (I_{eA} + I_{eB}) d^2}.$$

If  $r_A = r_B = r_{sc}$ , and  $I_e = I_{eA} = I_{eB}$  then Eqn. 34 becomes,

$$\text{Eqn. 35} \quad \frac{F_C}{F_J} = 2\sqrt{2}pe_0 \sqrt{\frac{q_{ion}}{m_{ion}}} \frac{r_{SC}^2 P^{3/2} e^{-d/l_d}}{I_e^3 d^2}.$$

For a formation of two spacecraft, we find that the  $F_C/F_J$  ratio is a function of the radii of the spacecraft, power supplied to the ion (electron) gun, the separation between the two spacecraft and the emission currents of both of them. Similar to the calculations for specific impulse, if we consider

formation of two identical spacecraft in GEO having same diameter of 1 m, charged to same high negative voltage  $V_{SC}$  and provided with same power  $P$  for each of them, they will each draw ion saturation current from the ambient plasma. So the (ion) emission current  $I_e$  will be also the same. Figure 14 shows the ratio of Coulomb to jet force assuming hydrogen ion emission in average GEO plasma.

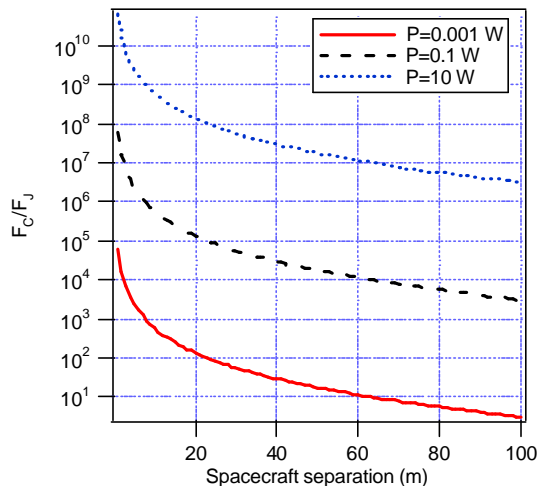


Figure 14. Comparison between induced Coulomb force and the momentum reaction of the emitted ion beam used to maintain the spacecraft charge for three power levels.

It can be seen from Figure 14 that for separations up to 100 m and system power greater than 1 mW the Coulomb force is considerably higher than the jet force. This implies two conclusions: 1) the Coulomb force is a wiser use of power than a mass-emitting EP thruster, and 2) the directional jet force will not be a significant perturbation to the Coulomb control system.

## 5. Conclusions

A new mode of potential spacecraft interaction was identified for closely spaced formations of vehicles in high Earth orbit. Using typical GEO conditions found in the literature, modeling and calculations were performed to show that Coulomb forces between vehicles may be as large as 1 mN for spacecraft 10 m apart, with electric dipole disturbance torques as high as 100  $\mu\text{N}\cdot\text{m}$  at the closest separations. For larger spacing most models showed the potential for forces of tens of  $\mu\text{N}$  and  $\mu\text{N}\cdot\text{m}$  torques persisting out to 50 m separation.

The Coulomb disturbance forces are commensurate with those expected from micropropulsion systems that will likely be used for formation maintenance. In an exploratory study, the possibility to purposefully exploit the Coulomb interactions as formation-keeping forces was investigated. The dynamic equations were formed within Hill's relative coordinate system for a collection of interacting vehicles. Analytic methods were developed to prove the existence of static equilibrium solutions using only Coulomb forces for propulsion within the swarm. Unique solution sets were found for one-dimensional three-spacecraft formations, a two-dimensional five-spacecraft formation, and a three-dimensional seven-spacecraft formation. Assuming 150-kg vehicles with physical dimensions of 1 meter, it was shown that kilovolt spacecraft potentials are sufficient to maintain most formations in a rigid geometry.

The Coulomb control system was evaluated in classical propulsion terms. It was shown that the specific impulse can be as high as  $10^{13}$  seconds. Due to the novel scaling of thrust, power, and specific impulse in this non-momentum thrust system, power levels as low as tens of milliwatts were shown sufficient to maintain the forces and to change their magnitude within a timescale of milliseconds.

## 6. Acknowledgements

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